

## Lower dimensional models in elasticity

With the purpose of fixing notation and nomenclature, we begin by quickly reviewing some fundamental notions in elasticity theory.<sup>1.1</sup> We then discuss dimension reduction in this context and its mathematical justification. We continue with a brief review of the literature where  $\Gamma$ -convergence is applied for this purpose, to conclude with an outline of the present work and some acknowledgements. Please refer to Appendix B for the notation used throughout this work.

### 1.1 Elasticity, in a rush

The objects of study are a three dimensional **body** identified with an open, bounded and Lipschitz set  $\Omega \subset \mathbb{R}^3$  and its **deformation**  $y: \Omega \rightarrow \mathbb{R}^3$  under external forces or boundary conditions. When deformations can be assumed to be very small it is more convenient to use instead **displacements**  $w: \Omega \rightarrow \mathbb{R}^3$ , defined by  $y(x) = x + w(x)$ . Throughout we employ so-called **Lagrangian coordinates**, i.e. we track the deformations of material points wrt. the fixed domain  $\Omega$ .<sup>1.2</sup>

Subject to external forces or boundary conditions, bodies deform. The fundamental assumption is that any deformation which is not a **rigid body motion** (the composition of a translation and a rotation) stores **elastic**

---

1.1. A thorough introduction to elasticity can be found in [Cia88], a gentle one from the perspective of differential geometry in [Cia05] and a deeper one in [MH94]. For a very good exposition of continuum mechanics with elasticity as an application see [TM05].

1.2. As opposed to the Eulerian description which instead tracks locations in space.

**energy** into the body which can be released after the extraneous conditions disappear and this release will bring the body back to its **reference configuration**  $\Omega$ , without inducing any permanent alteration. If this does not hold, that is, in case the properties of the body are changed after the forces disappear, one can have **viscoelastic** or **plastic** behaviour, but we will not concern ourselves with these at all. If the reference configuration has zero elastic energy, we speak of a **natural state**. The elastic energy can be computed as the integral over  $\Omega$  of a **stored energy density**  $W$ , which under mild assumptions turns out to be a function only of the position  $x \in \Omega$  and the **deformation gradient**  $\nabla y(x)$ . When this is the case we speak of a **hyperelastic** material. The function  $W$  expresses the relationship between **strains** (local elongations and compressions in each direction) and **stresses** (internal forces induced by the strains). By our fundamental assumption above,  $W$  is non-negative and vanishes for rigid motions, or  $W(x, \nabla y) = 0$  for all  $\nabla y \in \text{SO}(3)$ .

We model the strain by the change in metric induced by the map  $y$  in the body wrt. the flat metric, via the so-called **Green - St.Venant's tensor**  $E(y) = \frac{1}{2}(\nabla^\top y \nabla y - I)$ . In terms of displacements  $w = y - \text{id}$ , this is  $E(w) = \frac{1}{2}(\nabla^\top w + \nabla w + \nabla^\top w \nabla w)$ . Now we can characterise a **rigid motion** or **rigid body movement** as a deformation  $y$  such that  $E(y) = 0$ , i.e.  $\nabla^\top y \nabla y = I$ , since there is no change in the distance between deformed points. The set of all rigid motions consists of all maps  $x \mapsto Qx + c$  with  $Q \in \text{SO}(3)$ ,  $c \in \mathbb{R}^3$ . Under the assumptions that displacements are “infinitesimally smaller” than the characteristic dimensions of the body,  $E$  is approximated by the **linear strain tensor**  $e(w) := \nabla_s w = (\nabla^\top w + \nabla w)/2$  and one speaks of **geometrically linear** elasticity.

Assuming a smooth energy density and a small displacement gradient  $\|\nabla w\| \ll 1$ , one can linearise the energy around the identity:

$$\begin{aligned} W(\nabla y) &= W(I) + DW(I)[\nabla w] + \frac{1}{2}D^2W(I)[\nabla w, \nabla w] + h.o.t. \\ &\approx \frac{1}{2}D^2W(I)[\nabla w, \nabla w] \\ &=: \frac{1}{2}Q_3(\nabla w), \end{aligned}$$

where we used that  $W$  vanishes on rigid motions so, in particular  $W(I)$  and  $DW(I)$  are zero, and where  $Q_3$  is the **quadratic form of linear elas-**

**ticity**. In this setting we speak of **linearly elastic** materials. The form  $Q_3$  vanishes exactly over the set of **linearised rigid motions**<sup>1,3</sup>

$$\mathcal{R} := \{x \mapsto Rx + b : R \in \text{so}(3), b \in \mathbb{R}^3\} = \{x \mapsto r \times x + b : r, b \in \mathbb{R}^3\},$$

where  $\text{so}(3)$  is the space of antisymmetric matrices.

In order to define  $Q_3$  in terms of the gradients  $\nabla w$  one needs so-called **constitutive relations** between stresses and strains, which may take into account properties like **isotropy** (the body exhibits no “preferred direction” along which responses are different) and **homogeneity** (the body has the same behaviour at any material point  $x \in \Omega$ ). The symmetries arising in isotropic, homogeneous materials imply that  $Q_3$  has the form

$$Q_3(F) = \lambda \text{tr}^2 F + 2 \mu |F|^2$$

where  $F = \nabla w \in \mathbb{R}_{\text{sym}}^{3 \times 3}$  is a strain tensor and  $\lambda, \mu$  are the **Lamé constants** of the material.

There are several other couples of physically meaningful magnitudes related to these two constants, among which we mention **Young's modulus**  $E$  and **Poisson's ratio**  $\nu$  since we use them in the implementation of the discretisations.  $E$  is a measure of how the body extends or contracts in response to tensile or compressive stresses.  $\nu$  measures the tendency of materials to compress in directions perpendicular to the direction of elongation.<sup>1,4</sup>

---

1.3. In the setting of very small displacements, one must exclude symmetries (large displacements) from rigid motions, which means that the rotation matrices  $Q$  do not have the eigenvalue  $-1$  and the maps  $I + Q$  are invertible. Then we can define  $R := (I - Q)(I + Q)^{-1}$  and recover  $Q$  with **Cayley's transform**  $R \mapsto (I - R)(I + R)^{-1} = Q$ . This bijection allows the identification of matrices  $Q$  with matrices  $R$ , so we can focus on maps  $x \mapsto Rx + b$  with  $R \in \text{so}(3)$ . Additionally, each  $R$  is determined by just 3 coefficients, so there exists a vector  $r \in \mathbb{R}^3$  such that  $Rx + b = r \times x + b$ .

1.4.  $E$  is defined as the quotients of stresses over strains along each direction, which reduces to a number for isotropic materials. Since strains are dimensionless, it has units of pressure  $N/m^2$  or Pa, with typical values in the mega- and gigapascal range.  $\nu$  is the quotient of transverse strain to axial strain, with a sign, for each direction. Again, for isotropic materials this is only a number. Typical values range from 0 for materials with insignificant transversal expansion when compressed (e.g. cork) to 0.5 for incompressible ones (e.g. rubber), but materials have been designed beyond this range (*auxetic metamaterials*).